

Mathematical Formulation on Non-Newtonian Dusty Fluid over Vertical Stretching Sheet

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Two-phase flow is the interaction of two different matters, namely fluid and solid that contributes to an intriguing model in the boundary layer flow problem. It can be precisely defined as the binary mixture of the fluid and spherical particles that involve interaction between both phases on the flow region. The fascinating class of non-Newtonian fluid, particularly Casson, Williamson and Jeffrey models are further considered for fluid phase. The mathematical formulation on the respective models are presented and discussed. The adiabatic two phases is engaged throughout the entire flow process by means of no phase changes occurred. The governing equations for the proposed models are presented herein and need to undergo the boundary layer approximation as way of tackling its difficulty into a solvable form by using the order of magnitude analysis.

Keywords: Mathematical formulation, non-Newtonian fluid, dusty fluid, vertical stretching sheet

I. INTRODUCTION

The field of fluid dynamics has been an object of research for the past century that relates to the fluid motion with forces. As regards to fluid, the viscosity is one of its fundamental characteristic which referring to the resistance of fluid when flowing (Nakayama, 2018). There is, therefore, the used of the Newton's law of viscosity as a remarkable way to measure the viscosity that holds for the shear stress being proportional to the velocity gradient. Air, water and electrolyte are categorized as Newtonian fluid since they portrayed such behavior. However, for those fluids that opposed the law are recognized as non-Newtonian fluid (Partal and Franco, 2010), which are commonly exhibited by blood, paint, polymer and etc. This complex fluid is classified into the different kinds of models depending on their rheological behavior that can be represented by the constitutive equation, the relation of applied stresses to deformation. Generally, the model is of

the three broad categories which are time independent, time-dependent and viscoplastic fluids (Nguyen and Nguyen, 2012). Investigating non-Newtonian fluids has received consideration attention due to its applications in the mining industry, lubrication and biomedical flows (Khan et al., 2018). The studies of non-Newtonian fluid with various conditions have been investigated by several researchers (Kasim et al., 2012; Aurangzaib et al., 2013; Arifin et al., 2018; Kasim and Shafie, 2010).

In the case of two-phase flow, the involvement of two different states of matter is considered, which can occur either between solid, liquid and gas. The solid-liquid phase for instance, is referring to the fluid having dust particles which are important in packed beds, centrifugal separation of particles, sedimentation, environment pollution, and blood rheology (Butt et al., 2017). By drawing on the concept of two-phase flow, Saffman (1962) has been able to show the stability of dusty gas, while Marble (1963) identified four similarity parameters

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for gas-particle flow. In view of this, the flow problems concerning fluid and dust particles have led to a numerous studies under different elements, effects, as well as dimensions (Isa et al., 2016; Sulochana et al., 2016; Arifin et al., 2017).

In this paper, Casson, Williamson and Jeffrey fluids among the time-independent fluid model are selected, which primary concern to study for fluid-particle flow with the influences of aligned magnetic field and mixed convection over a vertical stretching sheet. The fluids were also examined by different authors (Hussanan et al., 2018; Nadeem et al., 2013; Hayat et al., 2011). Before proceeding to solve the two-phase flow problems, it is necessary to formulate the basic equations that are reliable. This paper first introduces the basic equations for each flow models in the vector form. Then, the boundary layer approximation is performed by applying the order of magnitude analysis. The resulting equations embodied the above arguments are finally expressed.

II. GOVERNING EQUATIONS

Turning now to the computation of flow problems, the modelling is carried out by solving its governing equations which are generally consist of continuity, momentum and energy equations. In addition, they are derived according to the three fundamental physical principles, which are mass is conserved, Newton's second law and energy is conserved (Anderson Jr, 2009). These laws are also applied for dust phase. The basic equation of momentum, in particular, is different for each flow case caused by the different constitutive relation of fluid model which are revealed in Table 1. Next, the formulation for two-phase flow is carried out by independently derived the governing equations for each phase (Siddiqi et al., 2015). Three fluid models that have been mentioned so far will be examined, all of which are embedded with the dust particles. Keeping in mind that for fluid phase, they share the same continuity and energy equations, however, the momentum equation is based on their respective constitutive relation.

Table 1. The constitutive relation for non-Newtonian fluid

Types of fluid	Constitutive equation
i. Casson fluid	$\boldsymbol{\tau} = \begin{cases} (\mu_B + \rho_y / \sqrt{2\pi}) 2e_{ij}, & \pi > \pi_c \\ (\mu_B + \rho_y / \sqrt{2\pi_c}) 2e_{ij}, & \pi < \pi_c, \end{cases} \quad (1)$ <p data-bbox="564 1346 1430 1507">where μ_B and ρ_y corresponds to the plastic dynamic viscosity of non-Newtonian fluid and fluid yield stress respectively, $\pi = e_{ij}e_{ij}$ while π_c being its critical value with e_{ij} is given by</p> $e_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right). \quad (2)$
ii. Williamson fluid	$\boldsymbol{\tau} = \mu_0 [1 + \Gamma \dot{\gamma}] \mathbf{A}_1, \quad \dot{\gamma} = \sqrt{(1/2) \text{trace}(\mathbf{A}_1^2)}, \quad (3)$ <p data-bbox="549 1738 1430 1839">where μ_0, Γ, and \mathbf{A}_1 are zero shear rate of limiting viscosities, time constant and first Rivlin-Erickson tensor that is equated to $\mathbf{A}_1 = \nabla \mathbf{V} + (\nabla \mathbf{V})^T$.</p>
iii. Jeffrey fluid	$\boldsymbol{\tau} = \frac{\mu}{1 + \lambda_2} \left[\mathbf{A}_1 + \lambda_1 \left(\frac{\partial \mathbf{A}_1}{\partial t} + \mathbf{V} \cdot \nabla \right) \mathbf{A}_1 \right], \quad (4)$ <p data-bbox="549 1977 1430 2063">where μ is dynamic viscosity, λ_1 and λ_2 corresponds to the relaxation time and ratio of relaxation to retardation, respectively.</p>

The following equations are thus expressed for fluid and dust phases as

Continuity equation:

$$\begin{aligned}\nabla \cdot \mathbf{V} &= 0, \\ \nabla \cdot \mathbf{V}_p &= 0,\end{aligned}\quad (5)$$

Momentum equation:

$$\begin{aligned}\rho \left(\frac{\partial \mathbf{V}}{\partial t} + \mathbf{V} \cdot \nabla \mathbf{V} \right) &= -\nabla p + \nabla \cdot \boldsymbol{\tau} + \mathbf{F}_p + \mathbf{F}_b, \\ \rho_p \left(\frac{\partial \mathbf{V}_p}{\partial t} + \mathbf{V}_p \cdot \nabla \mathbf{V}_p \right) &= -\mathbf{F}_p,\end{aligned}\quad (6)$$

Energy equation:

$$\begin{aligned}\rho c_p (\mathbf{V} \cdot \nabla) T &= k \nabla^2 T + Q_p, \\ \rho c_s (\mathbf{V} \cdot \nabla) T_p &= -Q_p,\end{aligned}\quad (7)$$

Here, \mathbf{V} , $\boldsymbol{\tau}$, \mathbf{F}_b , T , ρ , c_p and k signifies to velocity field, stress tensor, body force, temperature, density, specific heat and thermal conductivity for fluid phase. Meanwhile, \mathbf{V}_p , T_p , and c_s refers to vector field, temperature and specific heat for dust phase. The total fluid-particle interaction of force and thermal per unit volume are respectively referred to

$$\mathbf{F}_p = 6\pi n \mu r_p (\mathbf{V}_p - \mathbf{V}), \quad (8)$$

$$Q_p = 4\pi n r_p k (T_p - T). \quad (9)$$

III. MATHEMATICAL ANALYSIS

The boundary layer theory is originally introduced by Prandtl (1904) to reduce the complexity of governing equations into the approximate solutions in which the only significant terms are retained. In an attempt to approximate the governing equations, each term is estimated with the order of magnitude which has been reported by Schlichting (1974). By taking into account the above arguments, this approach is applied into Equations (5)-(7) to obtain the desired results in which the retaining terms are determined to be important. In the Cartesian coordinates, the steady two dimensional flow is considered in which the basic equations of fluid and dust phases are expressed in x and y directions with

independent of time. The flow configuration of this problem is illustrated in Figure 1.

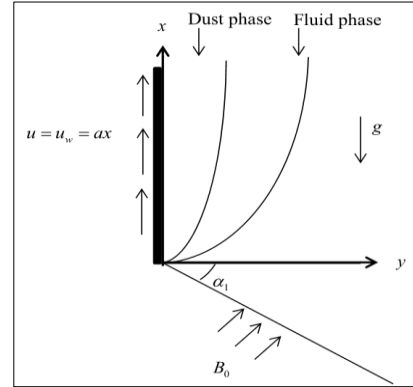


Figure 1: Flow configuration

To begin with, the terms contained in Equations (5)-(7) are estimated to be of order 1 and δ . Hence, the physical quantities contained in the respective equations are given as (Nadeem et al., 2013; Schlichting, 1974; Jaluria, 1980; Kasim, 2014; Kannan, 2001)

$$\begin{aligned}u &\sim 1, x \sim 1, v \sim \delta, y \sim \delta, \frac{\mu}{\rho} \sim \delta^2, \Gamma \sim \delta, \\ \sigma &\sim \frac{1}{\delta^2}, B_0^2 \sim \delta^2, \beta g (T - T_\infty) \sim 1, T \sim 1, \\ \alpha &\sim \delta^2, \lambda_1 \sim 1, \frac{\nu}{1 + \lambda_2} \sim \delta^2, \alpha \sim \delta^2 \\ u_p &\sim 1, v_p \sim \delta, \rho_p \sim \rho, \frac{1}{\tau_v} \sim 1, \frac{1}{\tau_T} \sim 1, T_p \sim 1\end{aligned}\quad (10)$$

A. Continuity Equation

In view of Equation (10), the continuity equation (5) is estimated and it can be seen from Table 2 that the first and second terms are of order 1 and they are considered to be retained.

Table 2. Order of magnitude for continuity equation

Term	Order of magnitude
$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$	$\frac{1}{1} \quad \frac{\delta}{\delta}$
$\frac{\partial u_p}{\partial x} + \frac{\partial v_p}{\partial y} = 0$	$\frac{1}{1} \quad \frac{\delta}{\delta}$

B. Momentum Equation

Substituting Equations (1)-(4) and (8) into (6) results in the momentum equation of fluid phase. Thus, the resulting equations for each fluid model are estimated by applying Equation (10) as shown in the following tables.

In the case of dusty Casson fluid, term $\partial^2 u / \partial x^2$ is negligible since it is of order δ^2 which is a very small quantity as presented in Table 3. Meanwhile, the remaining terms of order 1 are retained. From Table 4, the terms with order δ are maintained in y momentum equation. This is implying to the Prandtl's theory which stated that the inertia terms and viscous terms are required to have the same order of magnitude. However, the variations of y momentum in the fluid flow can be ignored stemming from all approximated terms are small quantity.

Summing up from the above elaborations, the approximation of y momentum equation for the other two fluid models will not be presented. Again, Schlichting (1974) has been shown that the equation of motion in y direction is completely discarded. For this reason, it is accurate enough to only estimate the momentum equation for x direction and the repeating terms as contained in dusty fluid are not included in the following tables.

Table 3. Order of magnitude for x momentum equation of dusty Casson fluid

Term	Order of magnitude
$u \frac{\partial u}{\partial x}$	$1 \frac{1}{1} = 1$
$v \frac{\partial u}{\partial y}$	$\delta \frac{1}{\delta} = 1$
$\frac{\mu_B}{\rho} \left(1 + \frac{1}{A}\right) \frac{\partial^2 u}{\partial x^2}$	$\delta^2 \frac{1}{1^2} = \delta^2$
$\frac{\mu_B}{\rho} \left(1 + \frac{1}{A}\right) \frac{\partial^2 u}{\partial y^2}$	$\delta^2 \frac{1}{\delta^2} = 1$
$\frac{\rho_p}{\rho} \frac{1}{\tau_v} u_p$	$\frac{\rho}{\rho} \frac{1}{1} = 1$
$\frac{\rho_p}{\rho} \frac{1}{\tau_v} u$	$\frac{\rho}{\rho} \frac{1}{1} = 1$
$\frac{\sigma}{\rho} B_0^2 \sin^2 \alpha_1 u$	$\frac{1}{\delta^2} \delta^2 = 1$

Table 4. Order of magnitude for y momentum equation of dusty Casson fluid

Term	Order of magnitude
$u \frac{\partial v}{\partial x}$	$1 \frac{\delta}{1} = \delta$
$v \frac{\partial v}{\partial y}$	$\delta \frac{1}{\delta} = \delta$
$\frac{\mu_B}{\rho} \left(1 + \frac{1}{A}\right) \frac{\partial^2 v}{\partial x^2}$	$\delta^2 \frac{\delta}{1^2} = \delta^3$
$\frac{\mu_B}{\rho} \left(1 + \frac{1}{A}\right) \frac{\partial^2 v}{\partial y^2}$	$\delta^2 \frac{\delta}{\delta^2} = \delta$
$\frac{\rho_p}{\rho} \frac{1}{\tau_v} v_p$	$\frac{\rho}{\rho} \frac{1}{1} \delta = \delta$
$\frac{\rho_p}{\rho} \frac{1}{\tau_v} v$	$\frac{\rho}{\rho} \frac{1}{1} \delta = \delta$
$\frac{\sigma}{\rho} B_0^2 \sin^2 \alpha_1 v$	$\frac{1}{\delta^2} \delta^2 = \delta$

Table 5. Order of magnitude for dusty Williamson fluid

Term	Order of magnitude
$\frac{\mu_0}{\rho} \frac{\partial^2 u}{\partial x^2}$	$\delta^2 \frac{1}{1^2} = \delta^2$
$\frac{\mu_0}{\rho} \frac{\partial^2 u}{\partial y^2}$	$\delta^2 \frac{1}{\delta^2} = 1$
$\frac{\mu_0}{\rho} \Gamma \left\{ \left(\frac{\partial u}{\partial x} \right)^2 + \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right)^2 \right\}^{\frac{1}{2}}$	$\delta^2 \delta \left\{ \frac{1}{\delta^2} \delta^2 \right\}^{\frac{1}{2}} = 1$
$\frac{\mu_0}{\rho} \Gamma \left\{ \left(\frac{\partial u}{\partial x} \right)^2 + \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right)^2 \right\}^{\frac{1}{2}} \left(\frac{\partial^2 u}{\partial y^2} \right)$	$\delta^2 \delta \left\{ \frac{1}{\delta^2} \right\}^{\frac{1}{2}} \left(\frac{1}{\delta^2} \right) = 1$

$$\frac{\mu_0}{\rho} \Gamma \frac{\partial}{\partial x} \left\{ \begin{array}{l} \left(\frac{\partial u}{\partial x} \right)^2 \\ + \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right)^2 \\ \left(\frac{\partial u}{\partial x} \right) \end{array} \right\}^{\frac{1}{2}} \delta^2 \delta \left\{ \begin{array}{l} 1 \\ \left(\frac{1}{\delta^2} \delta^2 \right) \end{array} \right\}^{\frac{1}{2}} \quad (1)$$

$$\frac{\mu_0}{\rho} \Gamma \frac{\partial}{\partial y} \left\{ \begin{array}{l} \left(\frac{\partial u}{\partial x} \right)^2 \\ + \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right)^2 \\ \frac{\partial u}{\partial y} \end{array} \right\}^{\frac{1}{2}} \frac{\delta^2 \delta}{\delta} \left\{ \begin{array}{l} \left(\frac{1^2}{1^2} \right) \\ \left(\frac{1}{\delta^2} \delta^2 \right) \end{array} \right\}^{\frac{1}{2}} \frac{1}{\delta}$$

$$\frac{\mu_0}{\rho} \Gamma \frac{\partial}{\partial y} \left\{ \begin{array}{l} \left(\frac{\partial u}{\partial x} \right)^2 \\ + \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right)^2 \\ \left(\frac{\partial v}{\partial x} \right) \end{array} \right\}^{\frac{1}{2}} \frac{\delta^2 \delta}{\delta} \left\{ \begin{array}{l} \left(\frac{1^2}{1^2} \right) \\ \left(\frac{1}{\delta^2} \delta^2 \right) \\ (\delta) \end{array} \right\}^{\frac{1}{2}}$$

$$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(\frac{\partial u}{\partial x} \frac{\partial^2 u}{\partial y^2} \right) \quad \delta^2 1 \left(\frac{1}{1} \frac{1}{\delta^2} \right) = 1$$

$$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(\frac{\partial v}{\partial x} \frac{\partial^2 u}{\partial x \partial y} \right) \quad \delta^2 1 \left(\frac{\delta}{1} \frac{1}{\delta} \right) = \delta^2$$

In the similar manner as dusty Casson fluid, the terms with order 1 from Tables 5 and 6 are remained. Therefore, according to Tables 3-6, the momentum equation for fluid phase of dusty Casson, dusty Williamson and dusty Jeffrey fluids are given by

$$u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} = \nu \left(1 + \frac{1}{A} \right) \frac{\partial^2 u}{\partial x^2} + \frac{\rho_p}{\rho \tau_v} (u_p - u) - \frac{\sigma}{\rho} B_0^2 \sin^2 \alpha_1 u + \beta g (T - T_\infty), \quad (12)$$

$$u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} = \nu \frac{\partial^2 u}{\partial y^2} + \sqrt{2\nu} \Gamma \frac{\partial u}{\partial y} \frac{\partial^2 u}{\partial y^2} + \frac{\rho_p}{\rho \tau_v} (u_p - u) - \frac{\sigma}{\rho} B_0^2 \sin^2 \alpha_1 u + \beta g (T - T_\infty), \quad (13)$$

Table 6. Order of magnitude for dusty Jeffrey fluid

Term	Order of magnitude
$\frac{\nu}{1 + \lambda_2} \frac{\partial^2 u}{\partial x^2}$	$\delta^2 \frac{1}{1} = \delta^2$
$\frac{\nu}{1 + \lambda_2} \frac{\partial^2 u}{\partial y^2}$	$\delta^2 \frac{1}{\delta^2} = 1$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(u \frac{\partial^3 u}{\partial x^3} \right)$	$\delta^2 1 \left(1 \frac{1}{1^3} \right) = \delta^2$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(u \frac{\partial^3 u}{\partial x \partial y^2} \right)$	$\delta^2 1 \left(1 \frac{1}{\delta^2} \right) = 1$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(v \frac{\partial^3 u}{\partial x^2 \partial y} \right)$	$\delta^2 1 \left(\delta \frac{1}{\delta} \right) = \delta^2$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(v \frac{\partial^3 u}{\partial y^3} \right)$	$\delta^2 1 \left(\delta \frac{1}{\delta^3} \right) = 1$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(\frac{\partial u}{\partial y} \frac{\partial^2 u}{\partial x \partial y} \right)$	$\delta^2 1 \left(\frac{1}{\delta} \frac{1}{\delta} \right) = 1$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(\frac{\partial u}{\partial y} \frac{\partial^2 v}{\partial x^2} \right)$	$\delta^2 \left(\frac{1}{\delta} \frac{\delta}{1^2} \right) = \delta^2$
$\frac{\nu}{1 + \lambda_2} \lambda_1 \left(\frac{\partial u}{\partial x} \frac{\partial^2 u}{\partial x^2} \right)$	$\delta^2 1 \left(\frac{1}{1} \frac{1}{1^2} \right) = \delta^2$

$$u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} = \frac{\nu}{1 + \lambda_2} \left[\begin{array}{l} \frac{\partial^2 u}{\partial y^2} \\ + \lambda_1 \left(\begin{array}{l} u \frac{\partial^3 u}{\partial x \partial y^2} + v \frac{\partial^3 u}{\partial y^3} \\ \frac{\partial u}{\partial y} \frac{\partial^2 u}{\partial x \partial y} - \frac{\partial u}{\partial x} \frac{\partial^2 u}{\partial y^2} \end{array} \right) \end{array} \right] + \frac{\rho_p}{\rho \tau_v} (u_p - u) - \frac{\sigma}{\rho} B_0^2 \sin^2 \alpha_1 u + \beta g (T - T_\infty), \quad (14)$$

In above equations, the Casson parameter is equated to $A = \mu_B \sqrt{2\pi} / \rho_y$, and the fluid viscosity ν can be defined according to each fluid model as $\nu = \mu_B / \rho$, $\nu = \mu / \rho$, $\nu = \mu_0 / \rho$. If $A \rightarrow \infty$, $\Gamma = 0$ and $\lambda_1 = \lambda_2 = 0$, then the fluid case of Newtonian fluid can be recovered. It is therefore important to note that the momentum equations (12)-(14) are under the Boussinesq approximation, that being applied to the case of flow with buoyancy force term. The same approaches have been discussed by dissimilar authors (Jaluria, 1980; Bejan and Kraus,

2003).

For dust phase, the momentum equation arises from inserting Equations (8) into (6) and it went on to use the boundary layer approximation that is treated similarly as in fluid phase by employing Equation (11). Tables 7 and 8 provide the analysis of order of magnitude for steady two dimensional of dust phase.

Table 7. Order of magnitude for x momentum equation of dust phase

Term	Order of magnitude
$u_p \frac{\partial u_p}{\partial x}$	$\frac{1}{1} = 1$
$v_p \frac{\partial u_p}{\partial y}$	$\delta \frac{1}{\delta} = 1$
$\frac{\rho_p}{\rho_p} \frac{1}{\tau_v} u_p$	$1 \cdot 1 = 1$
$\frac{\rho_p}{\rho_p} \frac{1}{\tau_v} u$	$1 \cdot 1 = 1$

Table 8. Order of magnitude for y momentum equation of dust phase

Term	Order of magnitude
$u_p \frac{\partial v_p}{\partial x}$	$\frac{\delta}{1} = \delta$
$v_p \frac{\partial v_p}{\partial y}$	$\delta \frac{\delta}{\delta} = \delta$
$\frac{\rho_p}{\rho_p} \frac{1}{\tau_v} v_p$	$1 \cdot \delta = \delta$
$\frac{\rho_p}{\rho_p} \frac{1}{\tau_v} v$	$1 \cdot \delta = \delta$

From the above tables, all terms in dust momentum equation remains unchanged since they are equally important. Despite that, the simplified form of this basic equation is expressed by dropping for y direction, hence

$$u_p \frac{\partial u_p}{\partial x} + v_p \frac{\partial u_p}{\partial y} = -\frac{1}{\tau_v} (u_p - u). \quad (15)$$

C. Energy Equation

The energy equation for fluid and dust phases are obtained by inserting Equations (9) into (7) and then the process of magnitude analysis is performed using

Equation (10) as provides in Tables 9 and 10.

Table 9. Order of magnitude for energy equation of fluid phase

Term	Order of magnitude
$u \frac{\partial T}{\partial x}$	$\frac{1}{1} = 1$
$v \frac{\partial T}{\partial y}$	$\delta \frac{1}{\delta} = 1$
$\alpha \frac{\partial^2 T}{\partial x^2}$	$\delta^2 \frac{1}{1^2} = \delta^2$
$\alpha \frac{\partial^2 T}{\partial y^2}$	$\delta^2 \frac{1}{\delta^2} = 1$
$\frac{\rho_p c_s}{\rho c_s} \frac{1}{\tau_T} T_p$	$1 \cdot 1 = 1$
$\frac{\rho_p c_s}{\rho c_s} \frac{1}{\tau_T} T$	$1 \cdot 1 = 1$

Table 10. Order of magnitude for energy equation of dust phase

Term	Order of magnitude
$u \frac{\partial T_p}{\partial x}$	$\frac{1}{1} = 1$
$v \frac{\partial T_p}{\partial y}$	$\delta \frac{1}{\delta} = 1$
$\frac{\rho_p c_s}{\rho c_s} \frac{1}{\tau_T} T_p$	$1 \cdot 1 = 1$
$\frac{\rho_p c_s}{\rho c_s} \frac{1}{\tau_T} T$	$1 \cdot 1 = 1$

According to Table 9, the term $\partial^2 T / \partial x^2$ is discarded from the fluid energy equation since it is estimated to be small quantity. Therefore, the energy equation for both phases can be written as follows that consist of the terms with order 1, we have

$$u \frac{\partial T}{\partial x} + v \frac{\partial T}{\partial y} = \alpha \frac{\partial^2 T}{\partial y^2} + \frac{\rho_p c_s}{\tau_T \rho c_p} (T_p - T), \quad (16)$$

$$u_p \frac{\partial T_p}{\partial x} + v_p \frac{\partial T_p}{\partial y} = -\frac{1}{\tau_T} (T_p - T). \quad (17)$$

IV. CONCLUSION

The indicating Equations (12)-(17) are now can be reduce into ordinary differential equations, which is less complex by using the appropriate similarity transformation which enable the numerical result to be computed.

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